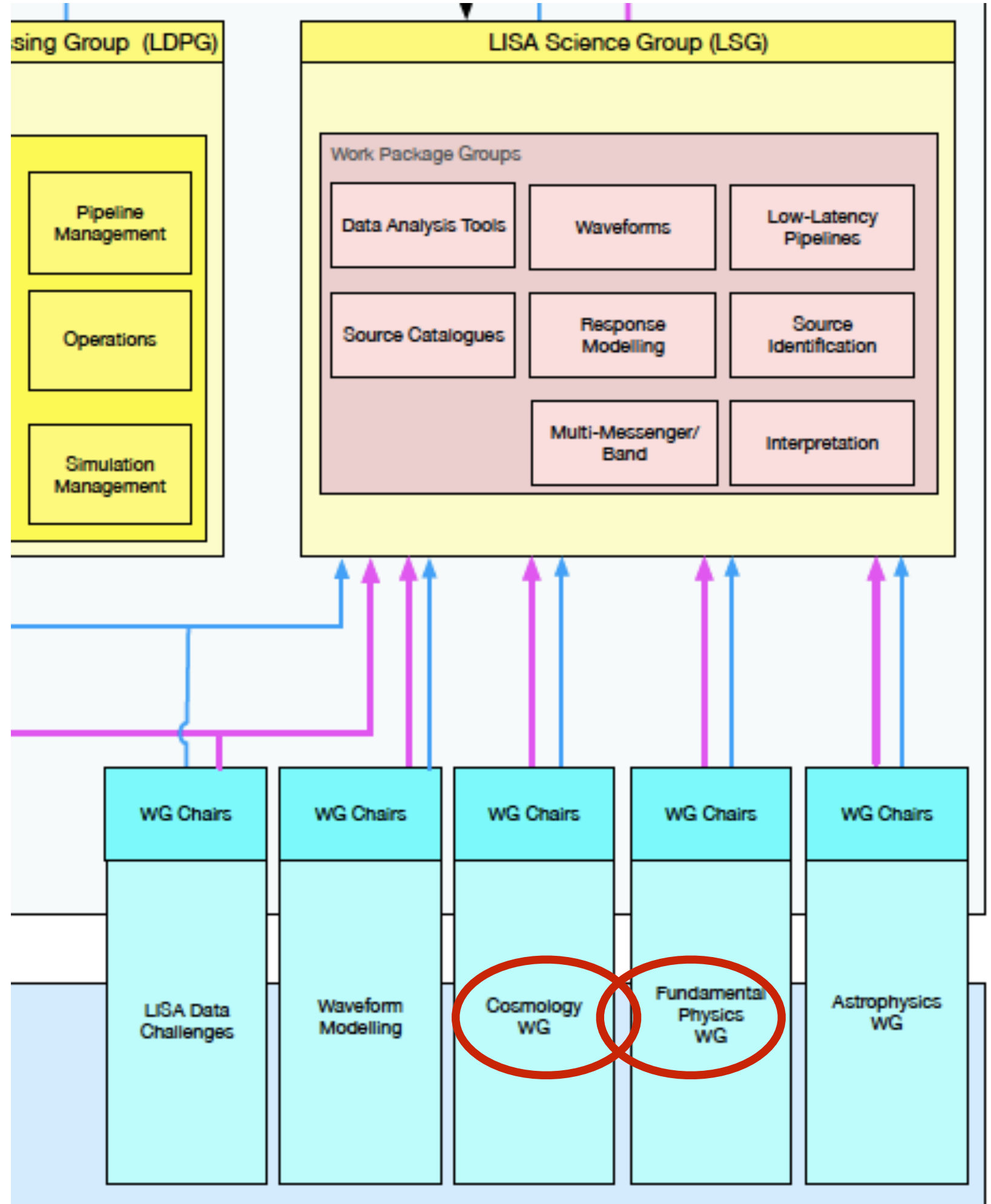
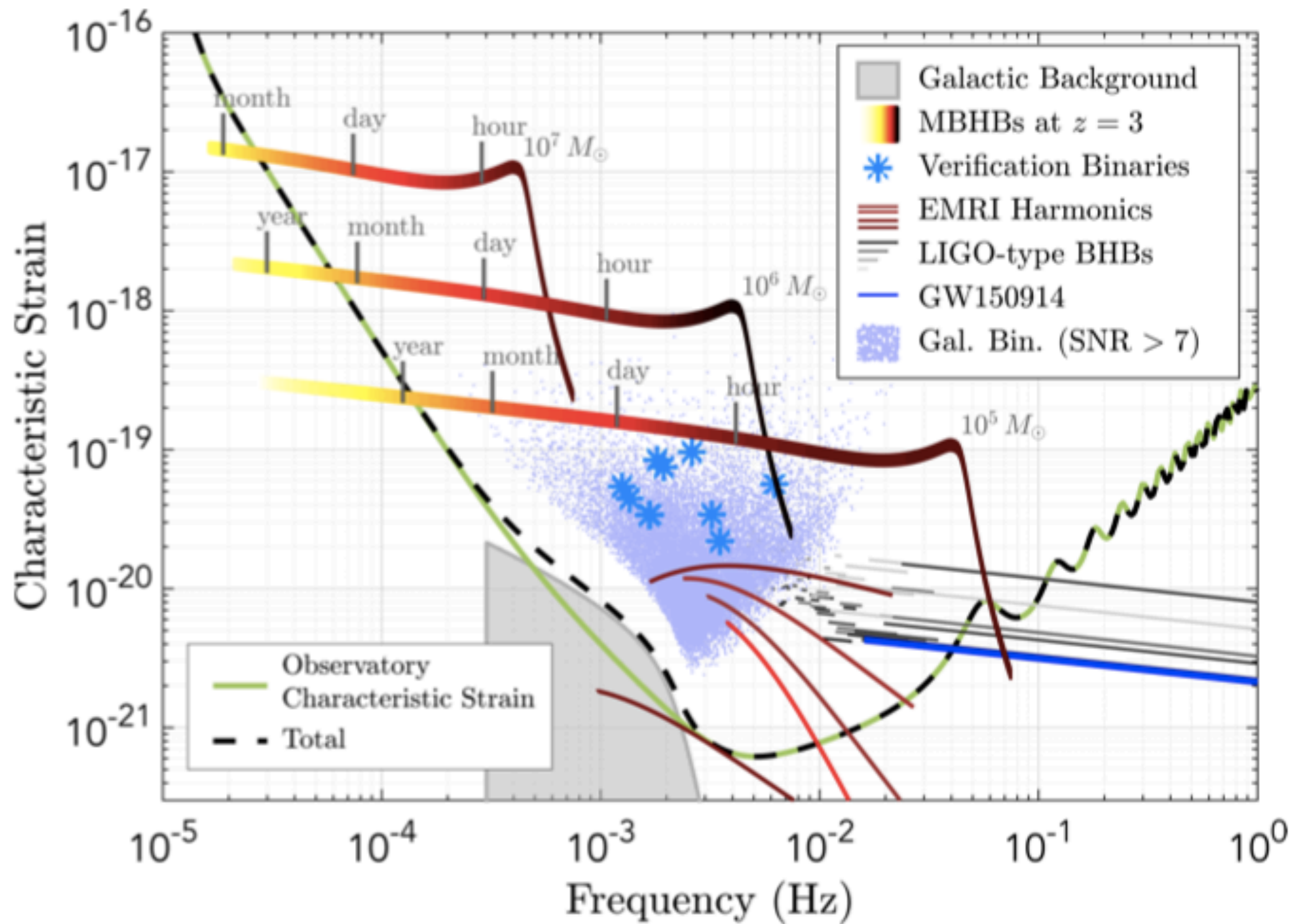


DISCUSSION ABOUT COSMOLOGY

WORKING
GROUPS:
interface
consortium
and
scientific
community



LISA sources



Cosmology with LISA

- GW emission from binaries
 - the stochastic GW background
-
- * *test of the early universe and high energy phenomena (mainly SGWB)*
 - * *test of the accelerated expansion of the universe (binaries)*
 - standard sirens: Lambda, Dark Energy, Modified Gravity **Michele**
 - effective field theory approach **Mark**
 - * *cross-correlation of GW sources and Large Scale Structure (both) **Alvise***
 - * *primordial black holes as dark matter (both) **Marco***

the stochastic GW background from primordial sources:
test of the early universe and high energy phenomena

amplification of vacuum fluctuations

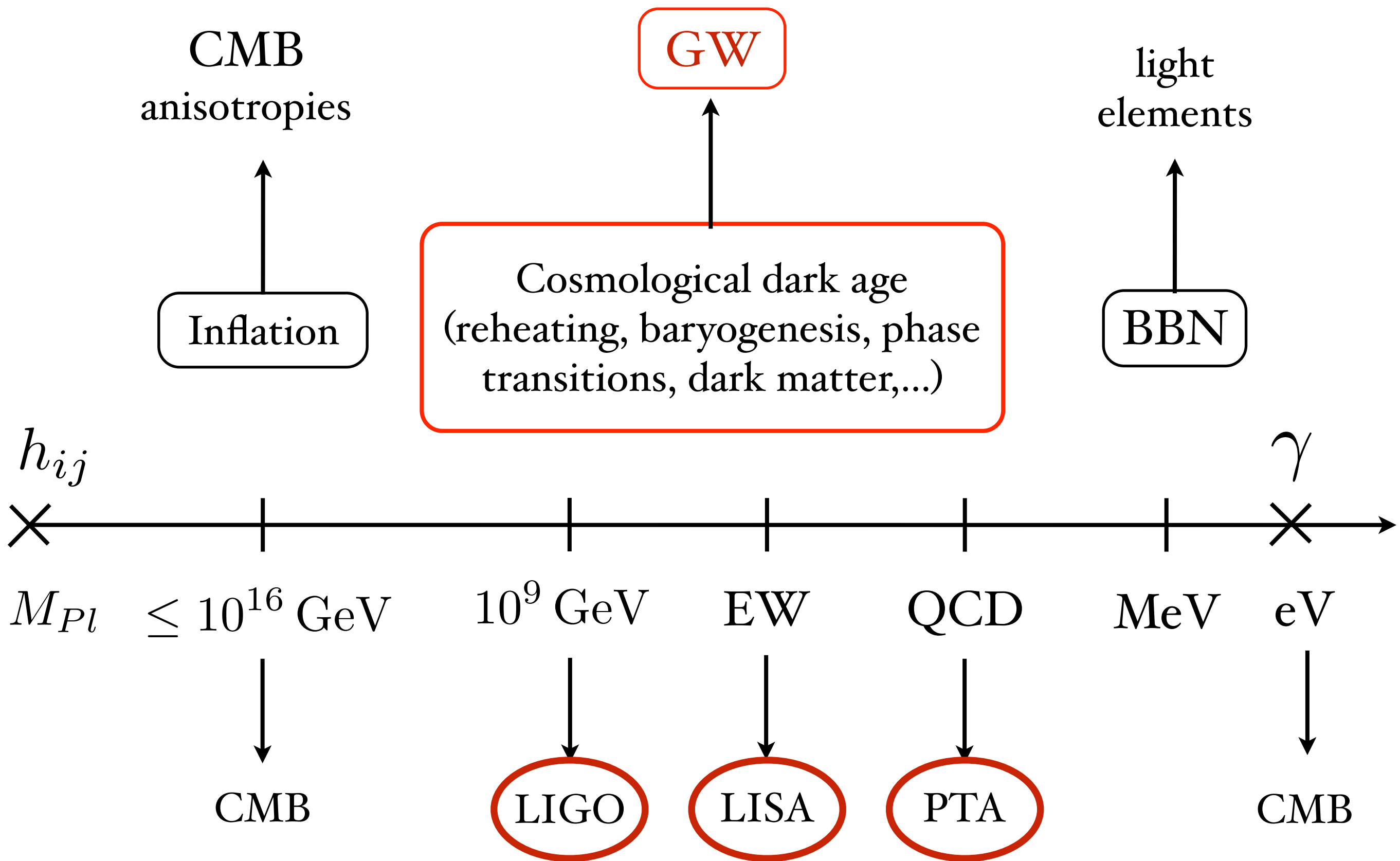
$$\ddot{h}_{ij} + 3H \dot{h}_{ij} + k^2 h_{ij} = 0$$

active source (tensor anisotropic stresses)

$$\ddot{h}_{ij} + 3H \dot{h}_{ij} + k^2 h_{ij} = 16\pi G \Pi_{ij}^{TT}$$

$$\epsilon_* = L_* H_* \quad f_c = f_* \frac{a_*}{a_0} = \frac{2 \cdot 10^{-5}}{\epsilon_*} \frac{T_*}{1 \text{ TeV}} \text{ Hz}$$


characteristic
frequency today



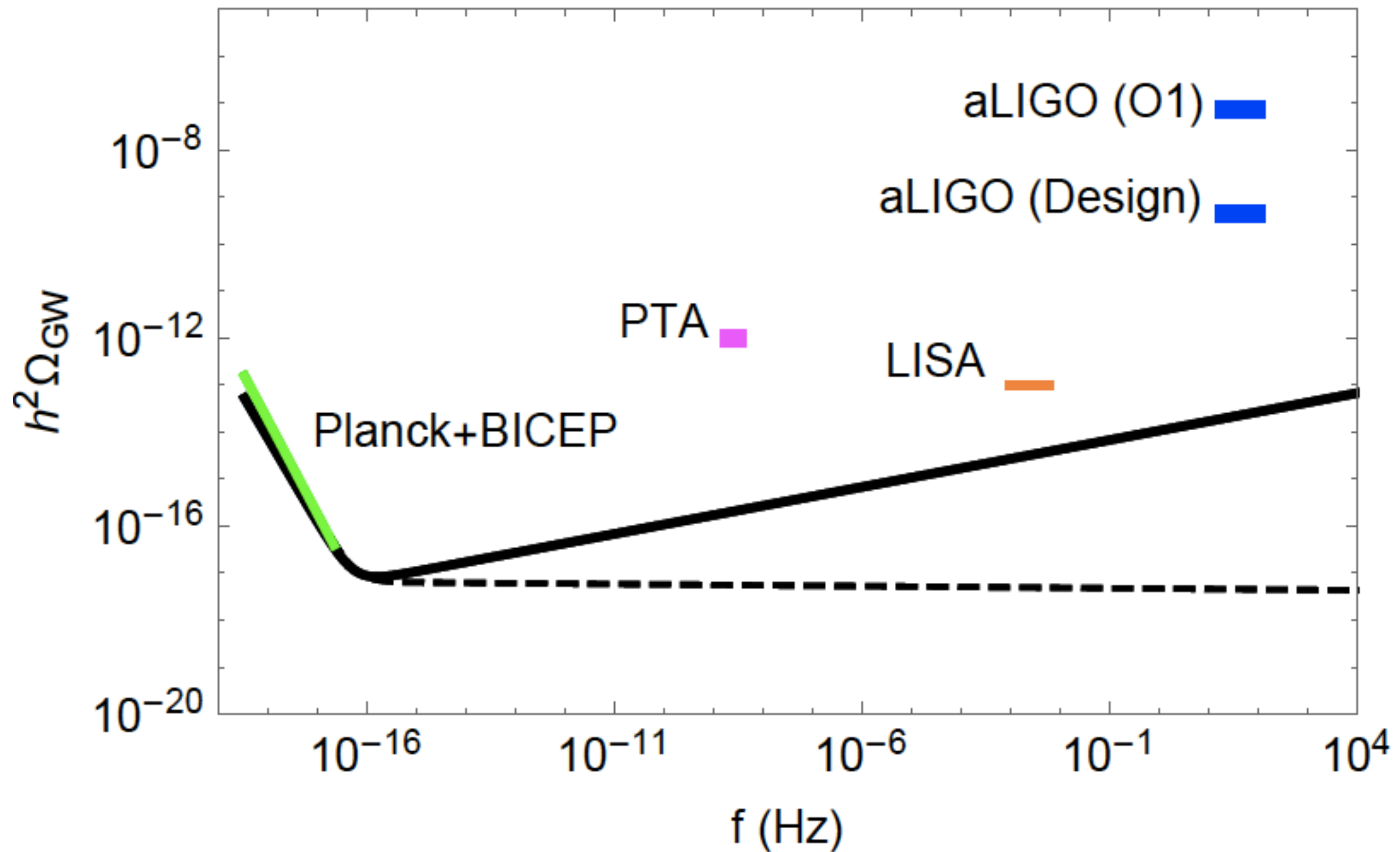
SGWB sources in the early universe : inflation-related

- irreducible SGWB from inflation
 - also sourced by second order scalar perturbations
- beyond the irreducible SGWB from inflation
 - particle production during inflation (scalar, gauge fields... coupled to the inflaton)
 - spectator fields
 - breaking symmetries (space-dependent inflaton, massive graviton)
 - modified gravity during inflation (massive GWs with $c \neq 1$)
 - primordial black holes
- alternatives to inflation
 - pre big-bang
 - cyclic/ekpyrotic
 - string gas cosmology
- preheating and non-perturbative phenomena
 - parametric amplification of bosons/fermions
 - symmetry breaking in hybrid inflation
 - decay of flat directions
 - oscillons

SGWB sources in the early universe : phase transition-related

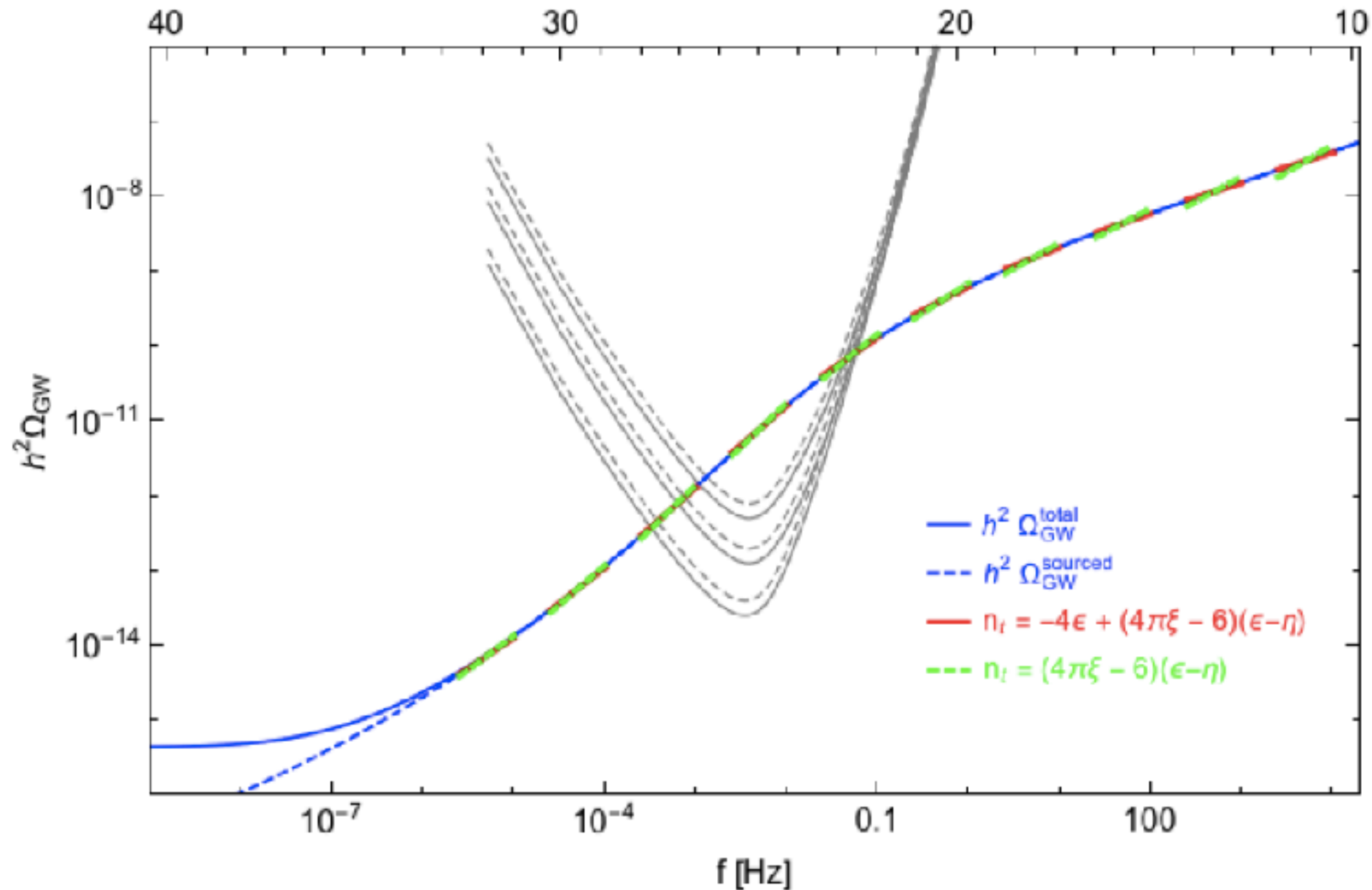
- first order phase transition
 - true vacuum bubble collision
 - sound waves
 - turbulence
- cosmic topological defects
 - irreducible SGWB from topological defect networks
 - decay of cosmic string loops
- thermal SGWB (standard model plasma)

SGWB sources in the early universe : inflation-related



just one example: inflaton-gauge field coupling

$$\Delta\mathcal{L} = -\frac{1}{4\Lambda}\phi F_{\mu\nu}\tilde{F}^{\mu\nu}$$



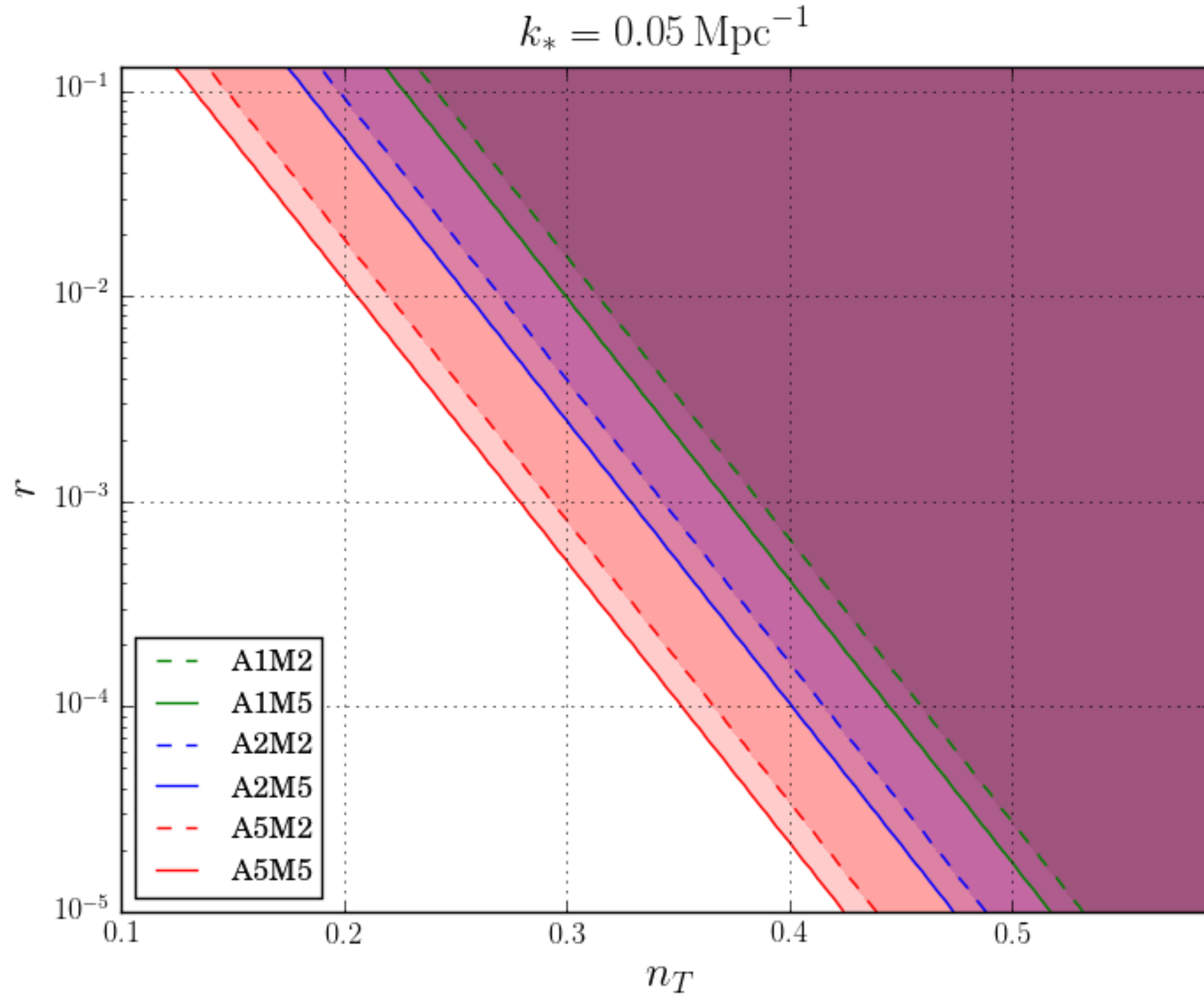
$$\Lambda = \frac{M_{Pl}}{35}$$

quadratic
inflaton
potential

OTHER SIGNATURES:
non-gaussianity, chirality

N. Bartolo et al, arXiv:1610.06481
N. Bartolo et al, arXiv:1806.02819

general constraints on (r, n_T) from LISA

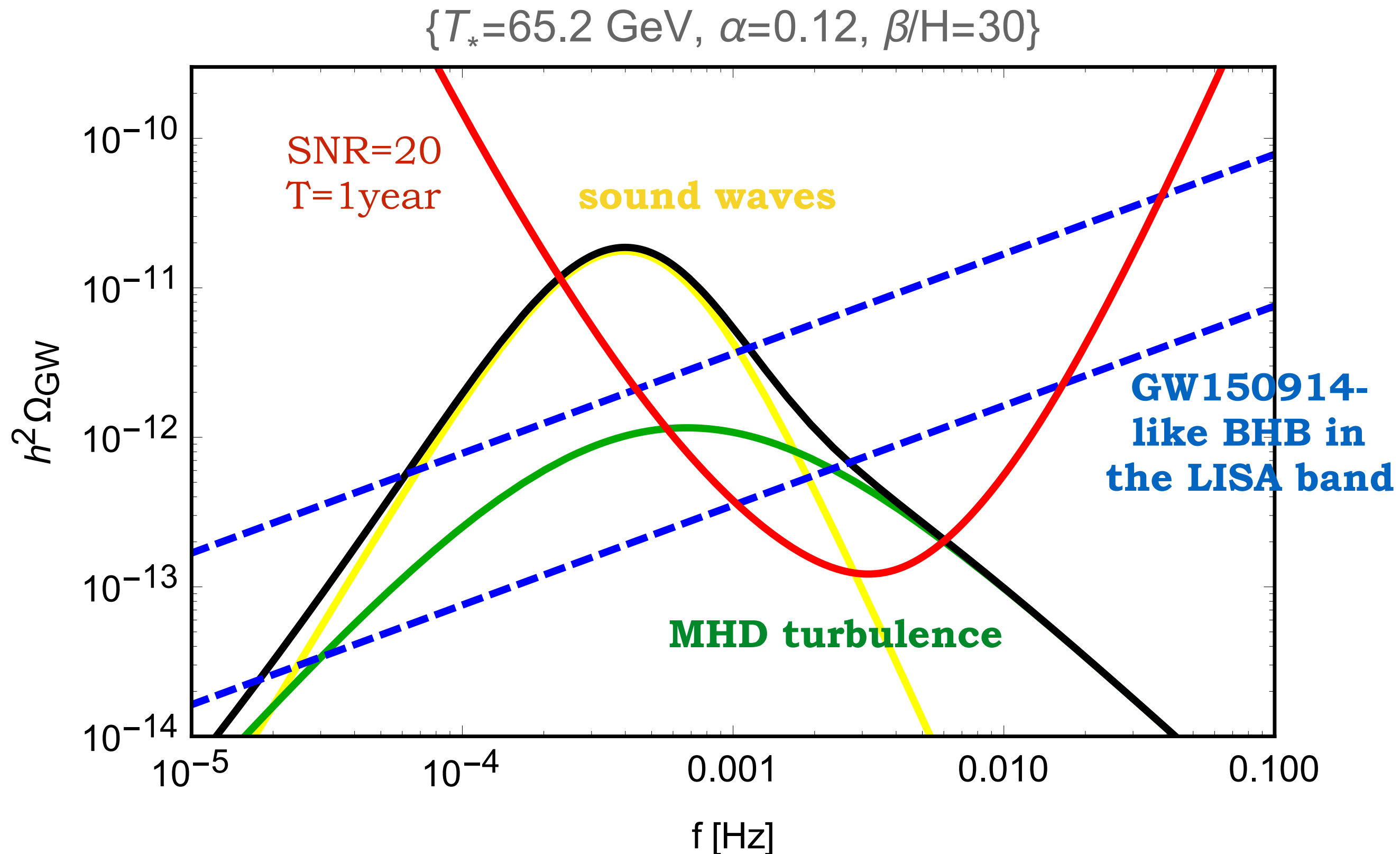


SGWB sources : phase transition-related first order PT

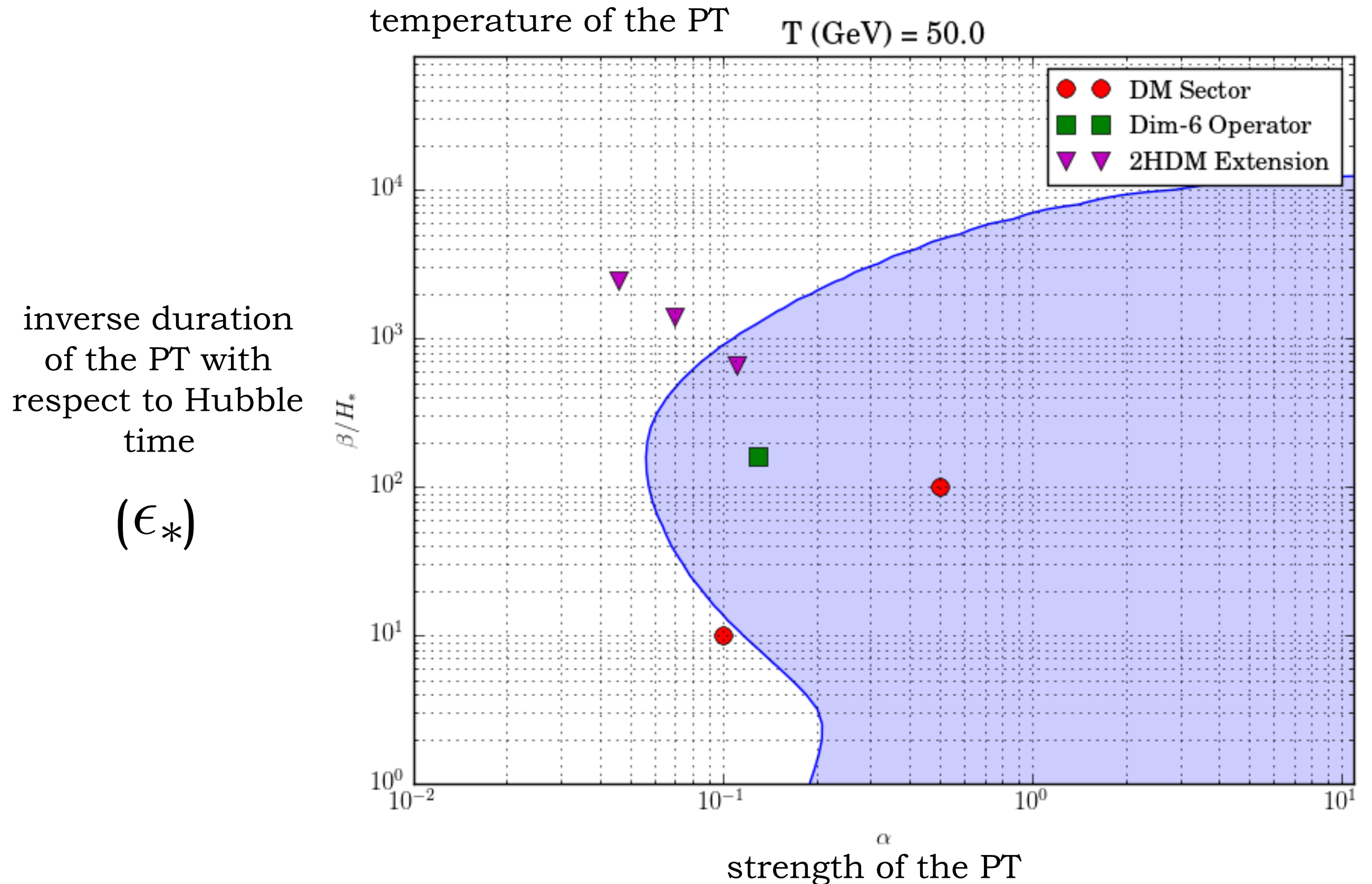
- LISA is sensitive to energy scales 10 GeV - 100 TeV
- LISA can probe the EWPT in BSM models ...
 - singlet extensions of MSSM
 - direct coupling of Higgs sector with scalars
 - SM plus dimension six operator
- ... and beyond the EWPT
 - Dark Matter sector
 - Warped extra dimensions

connections with baryon asymmetry, dark matter : LISA as a
complementary probe of BSM physics

Example of signal from FOPT + foreground

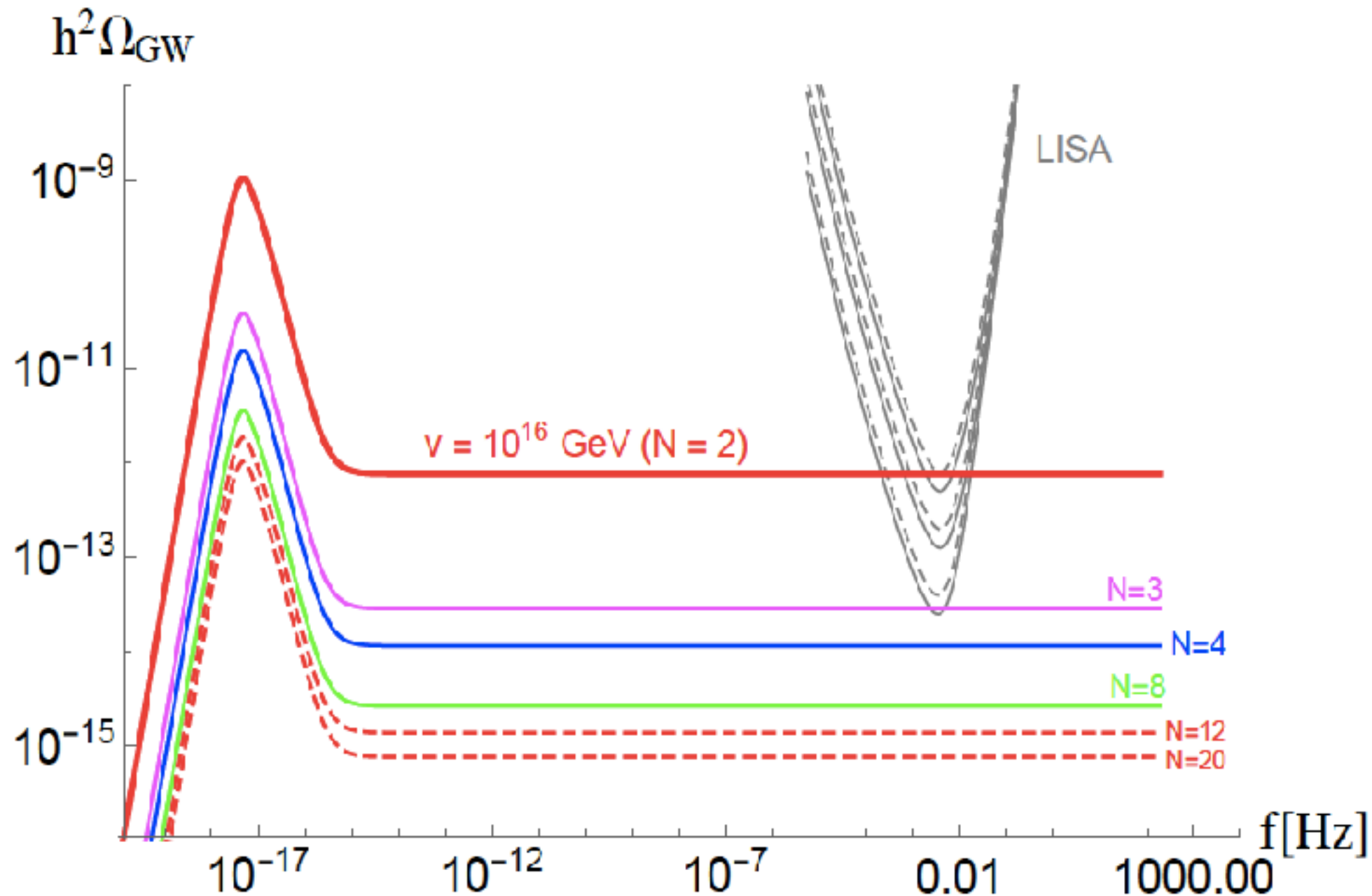


Example of detection prospects for FOPT at LISA



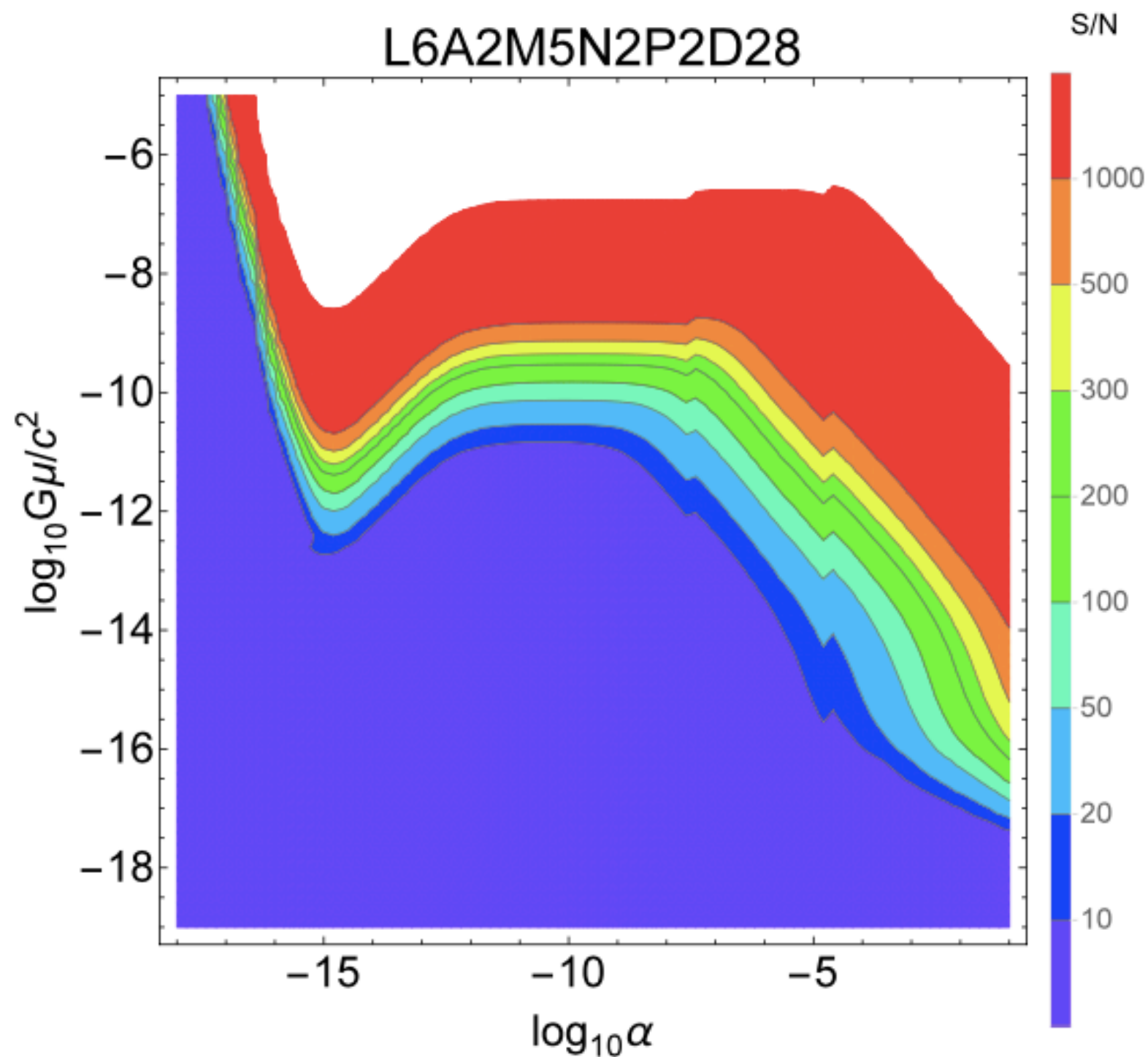
SGWB sources : phase transition-related topological defects

irreducible SGWB from a network of defects :
SSB of $O(N)$ global symmetry



Bounds on Nambu-Goto strings, loop size

S. Sanidas, LISA internal (old) $G\mu < 10^{-17}$



Current
NanoGRAV

$$G\mu < 1.3 \cdot 10^{-10}$$

future
CMB B-modes

$$G\mu < 10^{-9}$$

Future SKA

$$G\mu < 10^{-13}$$

Janssen et al 2015

Conclusions

- * LISA has a great potential to constrain early universe scenarios through the observation of a SGWB
- * inflationary models, EW symmetry breaking and beyond...
- * provided we can distinguish the signal from noise and astrophysical foregrounds
- * provided we can ascertain the origin of the SGWB once we measure it